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Fermi Gas assumptions:

1.) Non-interacting fermions

2.) Equi-distant single particle spacing $\rho(U) \propto \exp\left[2\sqrt{aU}\right] U^{3/2}$ generally successful

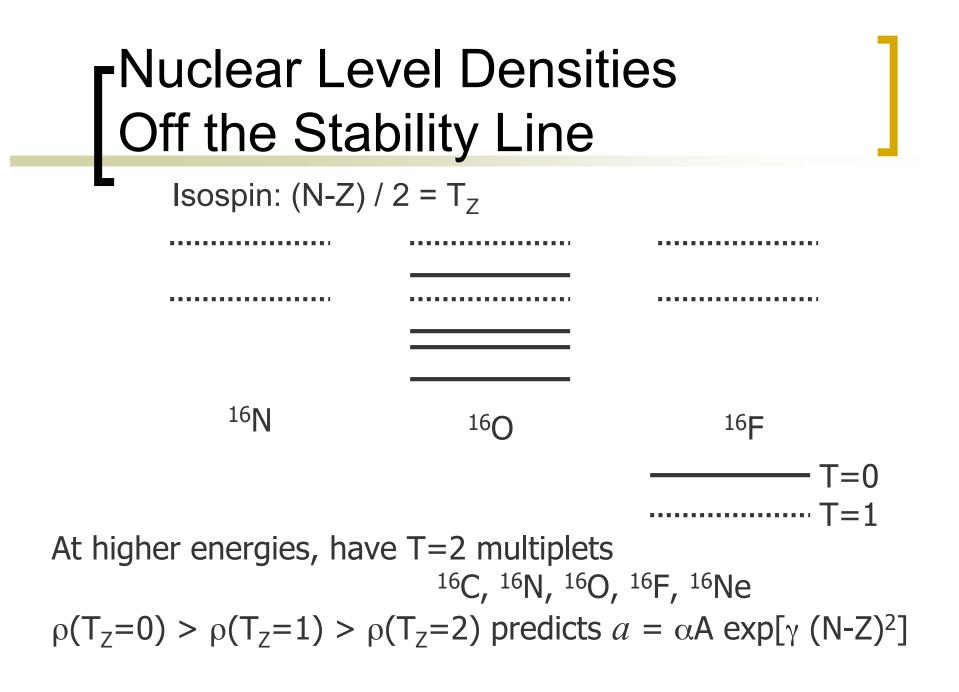
Most tests at U < 20 MeV for nuclei near the bottom of The valley of stability

Data: (1) Neutron resonances U ≈ 7 MeV near valley of stability (2) Evaporation spectra $U \approx 3$ to 15 MeV near valley of stability (3) Ericson fluctuations $U \approx 15$ to 24 MeV near valley of stability (4) Resolved levels $U \le 4$ to 5 MeV most points for nuclei in bottom of valley of stability

Study of Al-Quraishi, et al. $20 \le A \le 110$

Found $a = \alpha \operatorname{A} \exp[-\gamma (Z-Z_0)^2]$ $\alpha \approx 0.11 \qquad \gamma \approx 0.04$ Z is Z of nucleus Z_0 is Z of β stable nucleus of that Z

Reduction in *a* is negligible if $|Z-Z_0| \le 1$ Significant effect for $|Z-Z_0| = 2$ Large effect for $|Z-Z_0| \ge 3$



CONCLUSION

Only one analysis finds *a* lower off of stability line

Limited data is the central problem

The bulk of level density information comes from neutron resonances Example: $n + {}^{32}S \rightarrow {}^{33}S^*$

At low energy (neutrons) find ½⁺ levels at about 7 MeV of excitation

Not feasible for unstable targets

Only get density at one energy

Need σ (= $\langle J_Z^2 \rangle^{1/2}$) to get total level density

PREDICTIONS

A compound nucleus state must have a width which is narrow compared to single particle width.

This indicates compound levels will not be present once occupancy of unbound single particle states is substantial.

Assume we can use Boltzmann distribution as approximation to Fermi-Dirac distribution

Since

$$U = a\theta^2, \quad \theta = \sqrt{U/a}$$

θ is temperature*a* is level density parameterU is excitation energy

If occupancy of state at excitation energy B is less than 0.1

$$\exp[-B/\theta] = \exp[-B\sqrt{a/U}] \le 0.1$$
$$a \approx A/8$$

exp[-B
$$\sqrt{A/8U}$$
] \leq 0.1,
So, -B $\sqrt{A/8U} \leq$ -2.3
U_C \leq (AB² / 42.3)

Above this energy we will lose states

B	U	Α	В	U
(MeV)	(MeV)		(MeV)	(MeV)
8	30.3	200	8	303.
6	17.04	200	6	170.0
4	7.58	200	4	75.8
2	1.9	200	2	19
	(MeV) 8 6 4	(MeV) (MeV) 8 30.3 6 17.04 4 7.58	(MeV) (MeV) 8 30.3 200 6 17.04 200 4 7.58 200	(MeV) (MeV) (MeV) 8 30.3 200 8 6 17.04 200 6 4 7.58 200 4

Т

DRIP LINE SYSTEMATICS

А	Z _n	Ζ ₀	Z _p
20	7.2	9.5	11.3
40	12.4	18.4	21.7
60	18.3	27.0	31.4
80	26.7	36.0	40.6
100	30.5	43.2	50.5
150	48.4	62.2	69.5
200	58.3	80.0	88.1

 Z_n is Z of neutron-drip-line nucleus; Z_0 is Z in the valley of stability; and Z_p is Z of proton-drip-line nucleus

For nucleosynthesis processes, we will frequently have $B \sim 4$ MeV for proton or neutron

For A \sim 20 level density is substantially reduced by 10 MeV if B = 4

For A ~ 20 level density limit is \approx B if B = 2 MeV

Substantial reduction in compound resonances

Will also reduce level density for final nucleus in capture.

We also must consider parity.

In *fp* shell even-even nuclei have more levels of + parity at low U than – parity

Even-odd or odd-even have mostly negative parity states

> Thus, compound nucleus states not only have reduced ρ(U) but also suppress *s*-wave absorption because of parity mismatch

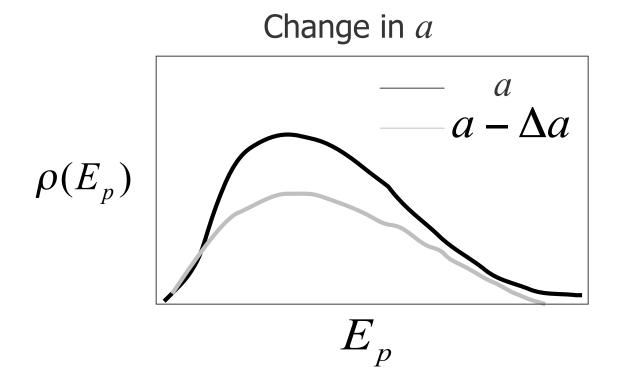
Experimental Tests

Measure:

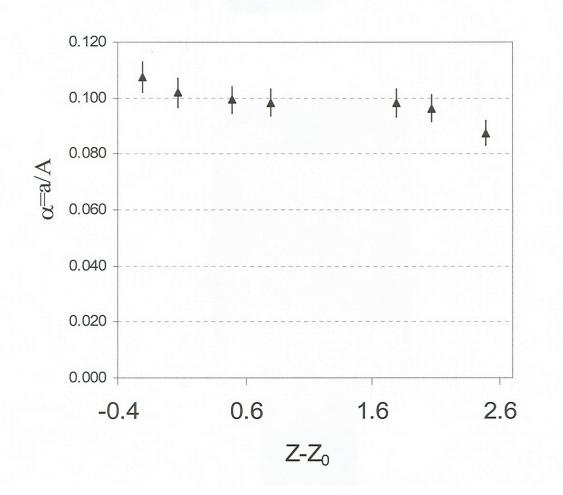
⁵⁵Mn(d,n)⁵⁶Fe

⁵⁸Fe(³He,p)⁶⁰Co
⁵⁸Fe(³He,α)⁵⁹Fe
⁵⁸Fe(³He,n)⁶⁰Ni

⁵⁸Ni(³He,p)⁶⁰Cu
⁵⁸Ni(³He,α)⁵⁷Ni
⁵⁸Ni(³He,n)⁶⁰Zn



Result: lower $a \rightarrow$ higher average E \rightarrow lower multiplicity If compound nucleus is proton-rich, Al-Quraishi term will further inhibit neutron decay



Span range of $|Z-Z_0|$ values up to ~ 2.5

Find evidence that a/A decreases with $|Z-Z_0|$ increasing

Currently calculating these effects:

Woods-Saxon basis Find single particle state energies and <u>widths</u> Compare ρ with all *sp* states with ρ including only bound <u>or</u> quasi-bound orbits

Expect reduction in *a* if $|Z-Z_0| \approx 2, 3$

As $|Z-Z_0|$ increases, new form with peak at 5 - 10 MeV may emerge (i.e. $\rho(20) \approx 0$)

Calculations and measurements underway:

- 1.) Hope to determine whether off of stability line *a* drops
- 2.) Is form $a = \alpha A \exp[-\gamma (Z-Z_0)^2]$ appropriate?
- 3.) As the drip line is approached, we may have to abandon Bethe form and switch to Gaussian

Heavy Ion Reactions

Can get to 30-40 MeV of excitation with lower pre-equilibrium component

Cannot use projectile with A about equal to target (quasi-fission)

Heavy Ion ReactionsRecent Argonne measurements $^{60}Ni + ^{92}Mo$ $^{60}Ni + ^{100}Mo$ got only 30-35% compound nuclear reactions $J_{contact}$ too highNot enough compound levels

Projectiles with A < half of target A are better Want compound nuclei with $|Z-Z_0| \ge 2$ to compare with $|Z-Z_0| \approx 0$

> REACTIONS ²⁴Mg + ⁵⁸Fe \rightarrow ⁸²Sr * ²⁴Mg + ⁵⁸Ni \rightarrow ⁸²Zr * ¹⁸O + ⁶⁴Ni \rightarrow ⁸²Kr *

Excitation energy $\approx 60 \text{ MeV}$ ${}^{82}\text{Kr} \text{ has } Z = Z_0$ ${}^{82}\text{Sr} \text{ has } Z = Z_0 + 2$ ${}^{82}\text{Zr} \text{ has } Z = Z_0 + 4$

Decay Fractions

	Rohr	Al-Quraishi
	0.96 n	0.93 n
⁸² Kr	0.04 p	0.06 p
	0.005 α	0.01 α
	0.67 n	0.61 n
⁸² Sr	0.25 p	0.33 p
	0.05 α	0.05 α
	0.02 d	0.01 d

⁸²Zr

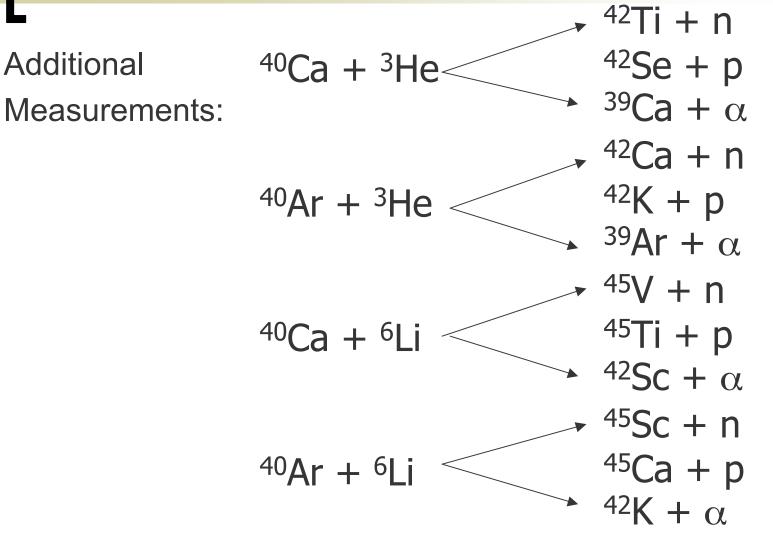
Decay Fractions

Rohr	Al-Quraishi
0.43 n	0.36 n
0.49 p	0.55 p
0.06 α	0.07 α
0.02 d	0.02 d

Rohr: Higher *a* for $|Z-Z_0| \ge 2$ Al-Quraishi: Lower *a* for $|Z-Z_0| \ge 2$

Get softer spectrum with Rohr More 4-6 particle emissions than Al-Quraishi

Compare: Rohr $a \propto \alpha A$ Al Quraishi $a \propto \alpha A \exp[\gamma (Z - Z_0)^2]$

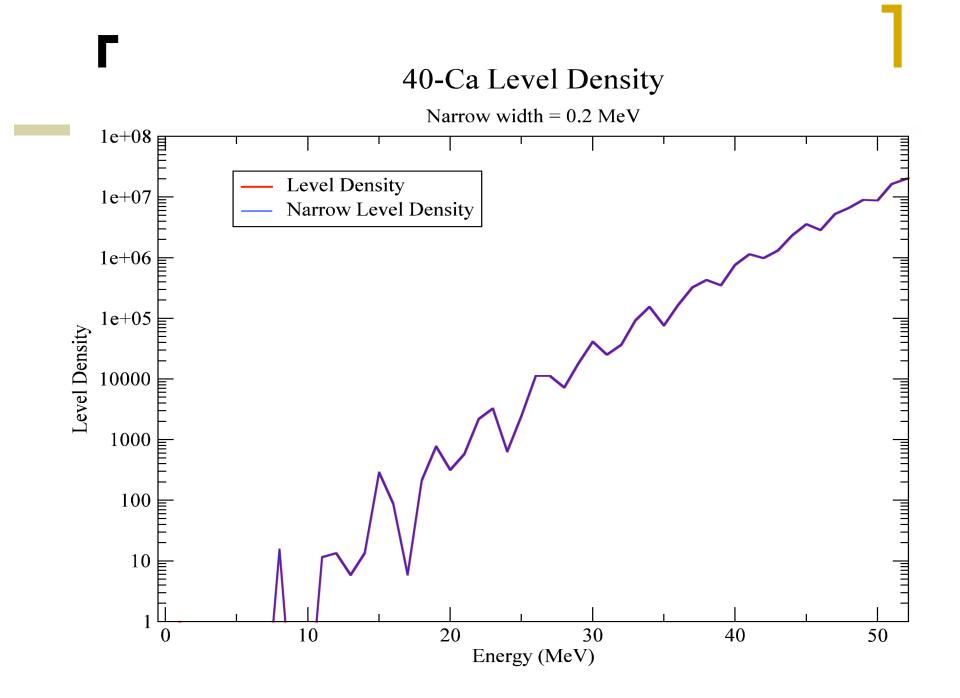


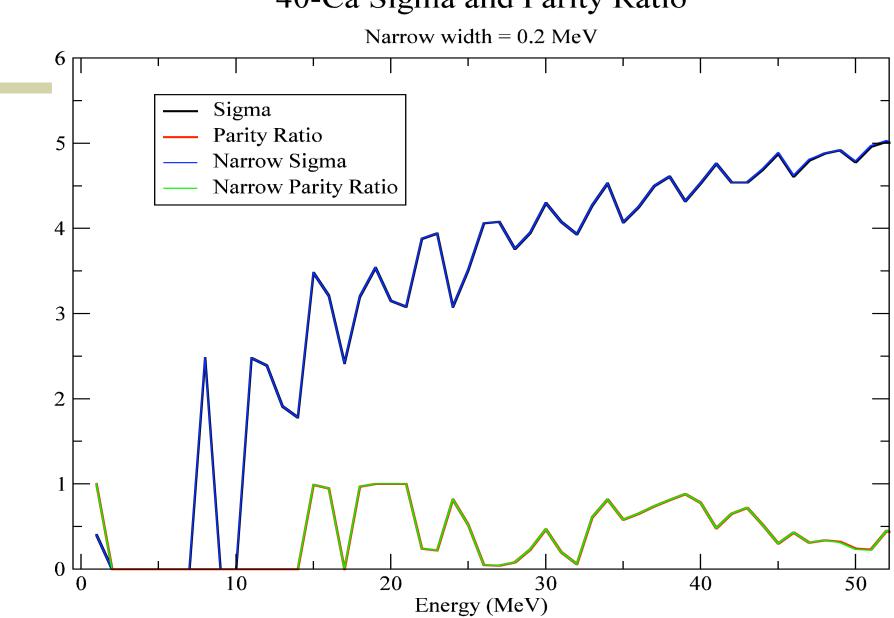
Calculations

- Solve for single particle energies in a single particle (Woods-Saxon) potential
- Compare level density including all single particle states with level density including only those with Γ < 500 keV
- Small effects for $Z \approx Z_0$

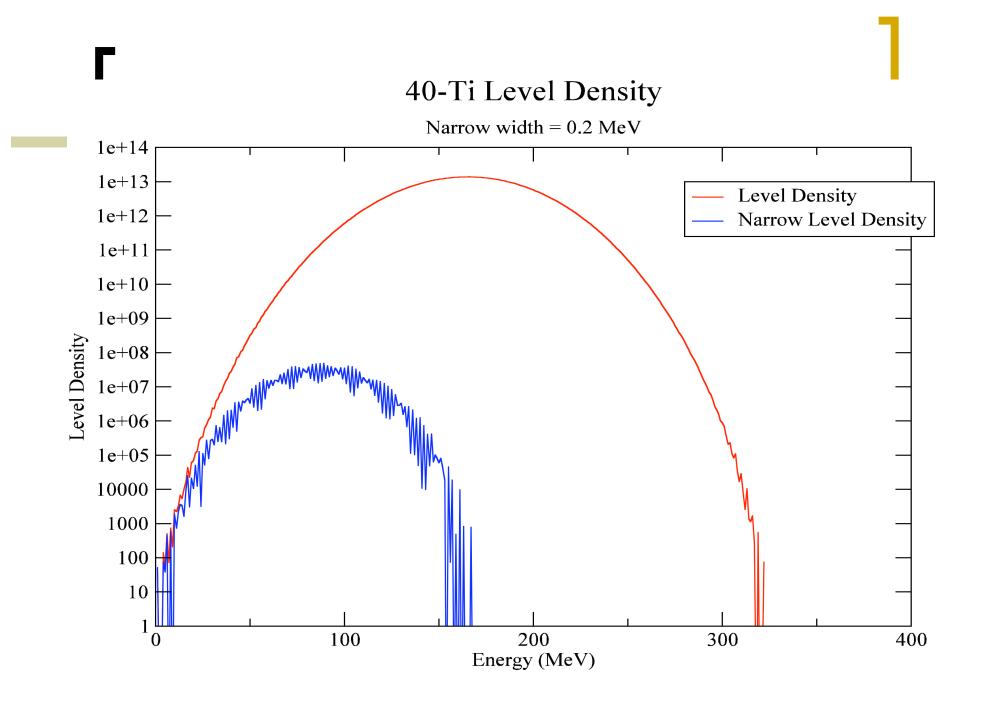
Large effects for $|Z-Z_0| \ge 4$

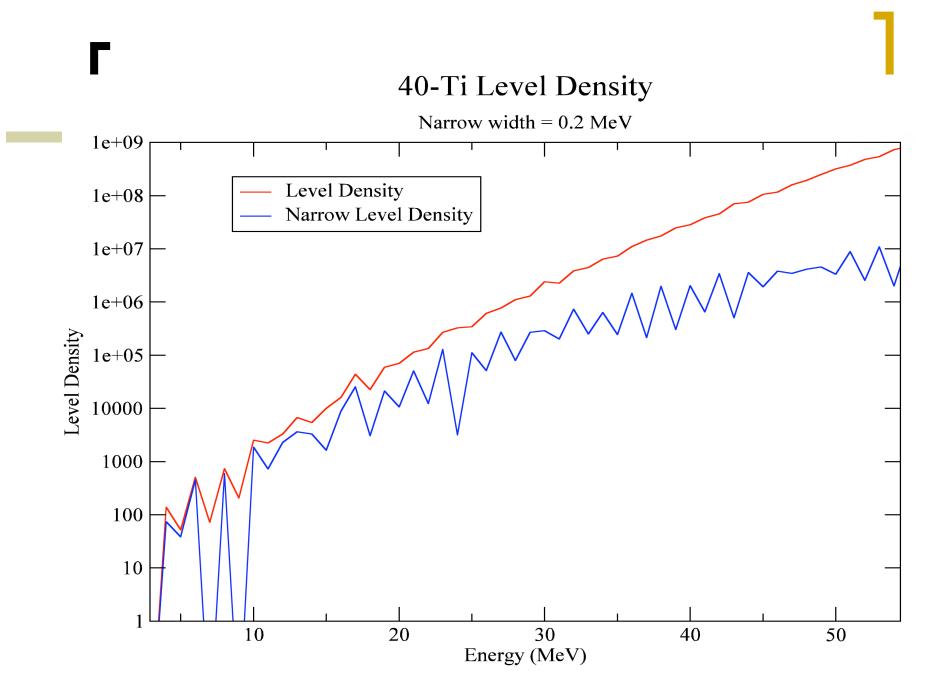
Looking at including two body effects in these calculations with moment method expansions



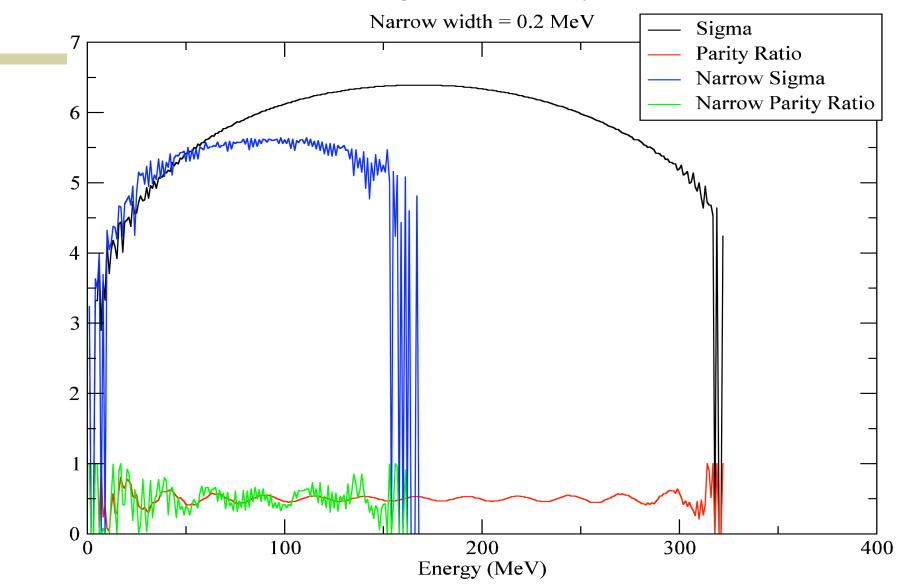


40-Ca Sigma and Parity Ratio





40-Ti Sigma and Parity Ratio



SUMMARY OF INITIAL CALCULATIONS

- A = $40:\rho \approx 0$ for Z = 12, Z = 24; *a* reduced by 15 to 20% for Z = 14 and Z = 18 (fit 0 to 20 MeV). Reduced by 30 to 40% for 0 to 100 MeV.
- A = 45: $\rho \approx 0$ for Z = 14, Z = 26; *a* reduced by 20-25% for Z = 16 and Z = 24 (fit 0 to 20 MeV). Reduced by 40% for 0 to 100 MeV.
- A = 50: $\rho \approx 0$ for Z = 16, Z = 28; *a* reduced by 25% for Z = 18 and Z = 26 (fit 0 to 20 MeV)

Other effects:

Increase fluctuations in ρ with U as $Z \rightarrow Z_n$, $Z \rightarrow Z_p$

Increase fluctuations in $\rho_{+}/(\rho_{+}+\rho_{-})$

Increase fluctuations in $\sigma(U)$

More likely that CN formation inhibited

Will examine A = 60, 70, 80, 100

Nuclear Level Densities Off the Stability Line CONCLUSIONS

Astrophysics has need for level densities off of the stability line for A \leq 100.

Data base in this region ($|Z-Z_0| \ge 2$) is poor

Model predicts that *a* decreases with $|Z-Z_0|$

Need more reaction data for nuclei in the region of $|Z-Z_0| \ge 2$

The End